

## Hot-spot emission properties in a warm plastic-shell implosion on OMEGA

W.L. Shang<sup>1</sup>, C. Stoeckl<sup>2</sup>, R. Betti<sup>2</sup>, S.P. Regan<sup>2</sup>, T.C. Sangster<sup>2</sup>, S.X. Hu<sup>2</sup>, A. Christopherson<sup>2</sup>, V. Gopalaswamy<sup>2</sup>, D. Cao<sup>2</sup>, W. Seka<sup>2</sup>, D.T. Michel<sup>2</sup>, A.K. Davis<sup>2</sup>, P.B. Radha<sup>2</sup>, F.J. Marshall<sup>2</sup>, R. Epstein<sup>2</sup>, A.A. Solodov<sup>2</sup>

<sup>1</sup> *Research Center of Laser Fusion, China Academy of Engineering Physics, Mianyang 621900 China*

<sup>2</sup> *Laboratory for Laser Energetics, University of Rochester, Rochester NY 14623*

A warm plastic-shell implosion was performed on the OMEGA laser system. The measured corona plasma evolution and shell trajectory in the acceleration phase are reasonably simulated by the one-dimensional LILAC simulation including the nonlocal and cross-beam energy transfer models. The results from analytical thin-shell model reproduce the time-dependent shell radius by LILAC simulation, and also the hot-spot x-ray emissivity profile at stagnation predicted by Spect3D. In the Spect3D simulations within a clean implosion, a “U”-shaped hot-spot radius evolution can be observed with the Kirkpatrick-Baez microscope response (the photon energy is from 4 to 8 keV). However, a fading away hot-spot radius evolution was measured in OMEGA warm plastic-shell implosion because of mixings. The distance from the measured hot-spot radius evolution shape to the “U” shape could be a new criterion for an experimental implosion performance. To recover the measured hot-spot x-ray emissivity profile at stagnation, a non-isobaric hot-spot model is built, and the normalized hot-spot temperature, density, and pressure profiles (normalized to the corresponding target-center values) are obtained.

### References:

- [1] V. N. Goncharov, T. C. Sangster, R. Betti, T. R. Boehly et al., Phys. Plasmas 21, 056315 (2014).
- [2] P. B. Radha, J. Delettrez, R. Epstein, V. Yu Glebov, R. Keck et al., Phys. Plasmas 9, 2208 (2002).